

1 **Data-driven transition path analysis yields a statistical understanding of**
2 **sudden stratospheric warming events in an idealized model**

3 Justin Finkel*

4 *Department of Earth, Atmospheric and Planetary Sciences, Massachusetts Institute of Technology*

5 Robert J. Webber

6 *Department of Computing and Mathematical Sciences, California Institute of Technology*

7 Edwin P. Gerber

8 *Courant Institute of Mathematical Sciences, New York University*

9 Dorian S. Abbot

10 *Department of the Geophysical Sciences, University of Chicago*

11 Jonathan Weare

12 *Courant Institute of Mathematical Sciences, New York University*

13 *Corresponding author: Justin Finkel, jfinkel@uchicago.edu

ABSTRACT

14 Atmospheric regime transitions are highly impactful as drivers of extreme weather events, but
15 pose two formidable modeling challenges: predicting the next event (weather forecasting), and
16 characterizing the average behavior over many events (the risk climatology). Each event has a
17 different duration and spatial structure, making it hard to define an objective “average event.” We
18 argue here that transition path theory (TPT), a framework from stochastic process theory, is an
19 appropriate tool for the task. We demonstrate TPT’s capacities on a wave-mean flow model of
20 sudden stratospheric warmings (SSWs) developed by Holton and Mass (1976), which is idealized
21 enough for transparent TPT analysis but complex enough to demonstrate computational scalability.
22 Whereas a recent article (Finkel et al. 2021) studied near-term SSW predictability, the present article
23 uses TPT to link predictability to long-term SSW frequency. This requires not only forecasting
24 forward in time from an initial condition, but also *backward in time* to assess the probability of the
25 initial conditions themselves. TPT enables one to condition the dynamics on the regime transition
26 occurring, and thus visualize its physical drivers with a vector field called the *reactive current*.
27 The reactive current shows that before an SSW, dissipation and stochastic forcing drive a slow
28 decay of vortex strength at lower altitudes. The response of upper-level winds is late and sudden,
29 occurring only after the transition is almost complete from a probabilistic point of view. This case
30 study demonstrates that TPT quantities, visualized in a space of physically meaningful variables,
31 can help to understand the dynamics of regime transitions.

32 1. Introduction

33 Many features of the atmosphere-ocean system’s large-scale variability can be understood, to
34 some extent, as transitions between qualitatively different regimes. Examples include blocking,
35 monsoons, El Niño, and Sudden Stratospheric Warming events (SSWs, the subject of this paper),
36 all of which are associated with extreme weather. From a scientific perspective, regime transitions
37 are handles by which to probe the climate’s nonlinear, non-equilibrium dynamics. They expose
38 novel physics and push us to qualitatively expand our physical understanding. From a human
39 perspective, these relatively rare anomalies pose major societal challenges (Lesk et al. 2016; Kron
40 et al. 2019), especially with a changing climate and increasing reliance on weather-susceptible
41 infrastructure (e.g., Mann et al. 2017; Frame et al. 2020).

42 Regime transitions are used as benchmarks for model development across the hierarchy, from
43 state-of-the-art Earth system models with billions of variables (e.g., Stephenson et al. 2008;
44 Lengaigne and Vecchi 2010; Vitart and Robertson 2018) to conceptual low-order models with
45 fewer than ten variables (e.g., Charney and DeVore 1979; Timmermann et al. 2003; Ruzmaikin
46 et al. 2003; Crommelin et al. 2004; Thual et al. 2016). In Finkel et al. (2021), we addressed
47 near term forecasting of regime transitions in the context of an idealized sudden stratospheric
48 warming (SSW) model due to Holton and Mass (1976), which possesses two metastable states: a
49 strong-vortex regime A , and a weak-vortex regime B .

50 The present paper’s main contribution is to address aspects of the long-term climate statistics
51 of SSW events: how often do they occur, what are their typical development pathways, and how
52 variable are those pathways between events? We will use the framework of transition path theory
53 (TPT; E and Vanden-Eijnden 2006), which offers a concise set of quantities to answer these
54 questions. An SSW event is represented as a *transition path* from A to B . The main quantity of

55 interest will be the *reactive current* \mathbf{J}_{AB} , defined in section 3, which specifies the flow of probability
56 density through state space *conditioned on an $A \rightarrow B$ transition event being underway*. To properly
57 implement that conditional statement, we will need two auxiliary quantities. First, the *forward*
58 *committor* $q_B^+(\mathbf{x})$ gives the probability that the system, initialized in a state \mathbf{x} , next reaches B before
59 A . This is a measure of progress toward SSW. Second, the *backward committor* $q_A^-(\mathbf{x})$ gives the
60 probability, looking backward in time, that the system visited A more recently than B .

61 The forward committor itself was a primary focus of Finkel et al. (2021), where we pursued
62 forecasting as a main objective. Committor probabilities are generally gaining traction as a metric
63 for weather prediction; see Tantet et al. (2015) for an application to atmospheric blocking, Lee et al.
64 (2018) for an application to tropical cyclone downscaling, Lucente et al. (2022) for an application to
65 El Niño, and Miloshevich et al. (2022) for a very recent application to heat waves. However, in the
66 present paper we are pursuing climatological statistics rather than forecasting probabilities, using
67 the committor only as an intermediate calculation for the reactive current, which characterizes the
68 full transition process from A to B rather than its “forward half” from \mathbf{x} to B .

69 Some previous studies (Crommelin 2003; Tantet et al. 2015) have visualized what are essentially
70 reactive currents for blocking events in an observable subspace of leading EOFs. However, these
71 studies were not couched in the language of TPT, a formalism that brings more quantitative results.
72 Namely, the reactive current \mathbf{J}_{AB} provides a direct estimate of the SSW rate, decomposing it over a
73 continuous probability distribution of pathways. Formal TPT has not yet been widely taken up by
74 the atmosphere-ocean science community, besides a few exceptions (Finkel et al. 2020; Miron et al.
75 2021, 2022). Part of our goal here is to encourage a common quantitative language for discussing
76 regime transitions, which could help to organize several existing lines of research.

77 \mathbf{J}_{AB} , like q_B^+ , can be expressed as a function of any observable subspace for visual exploration, with
78 the complementary subspace treated as random variables. It is most enlightening to use observables

79 with concrete physical meaning. A recent article Miloshevich et al. (2022) exploited this property to
80 interpret a neural-network-learned committor for heat waves in terms of geopotential height and soil
81 moisture, thus quantifying their predictive power at various lead times. In Finkel et al. (2021), we
82 visualized the committor and expected lead time in a two-dimensional subspace consisting of zonal
83 wind U , an index for polar vortex strength, and vertically integrated heat flux (IHF), which roughly
84 measures the amplitude and phase tilt of vortex-disrupting planetary waves. Here we continue to
85 use those coordinates, but also introduce a new subspace based on the zonal-mean meridional
86 potential vorticity (PV) gradient and eddy enstrophy. These two quantities obey a conservation law
87 in the absence of dissipation and stochastic forcing, a slight variation of the Eliassen-Palm relation.
88 This allows us to diagnose more precisely the crucial roles of dissipation and stochastic forcing
89 in driving the transition process, an important step toward understanding their causal relationship.
90 Other kinds of atmospheric regime transitions will have different relevant physical diagnostics, any
91 of which can be seen as an independent variable for the committor function and reactive current.

92 This paper is organized as follows. In section 2 we recapitulate the dynamical model. In section
93 3 we visualize the evolution of SSW events using the probability current, and introduce the key
94 quantities for TPT—committors, densities, and currents—along with a brief summary of the method
95 to compute them, which is more thoroughly explained in the supplementary document. In section
96 4, we use reactive current to construct a composite SSW evolution, and compare this to the standard
97 composite method. In section 5, we change coordinates to better examine the dynamics of SSW
98 events. We assess future directions and conclude in section 6.

99 **2. The dynamical SSW model**

100 We use exactly the same model as in Finkel et al. (2021), which is presented here for completeness.

101 *a. Model specification*

102 Holton and Mass (1976) developed a minimal model for the variability of the winter stratospheric
 103 polar vortex, capturing the wave-mean flow interactions behind sudden stratospheric warming
 104 events. The model's prognostic variables consist of a zonally averaged zonal wind $\bar{u}(y, z, t)$ and a
 105 perturbation geostrophic streamfunction $\psi'(x, y, z, t)$ on a β -plane channel with a central latitude
 106 of $\theta = 60^\circ\text{N}$, a meridional extent of 60° , and a height of 70 km, with the coordinate z ranging from
 107 0 at the bottom of the domain (the tropopause) to 70 km at the top of the domain. \bar{u} and ψ' are
 108 projected onto a single zonal wavenumber $k = 2/(a \cos \theta)$ and a meridional wavenumber $\ell = 3/a$:

$$\bar{u}(y, z, t) = U(z, t) \sin(\ell y) \quad (1)$$

$$\psi'(x, y, z, t) = \text{Re}\{\Psi(z, t)e^{ikx}\}e^{z/2H} \sin(\ell y), \quad (2)$$

109 where $a = 6370$ km is the radius of Earth, and $H = 7$ km is the scale height. U (the mean flow)
 110 and Ψ (a complex-valued wave amplitude) evolve according to the projected primitive equations
 111 and the linearized quasi-geostrophic potential vorticity (QGPV) equation. A non-dimensionalized
 112 version of the equations is as follows, rearranged slightly from Finkel et al. (2021). The mean flow
 113 $U(z, t)$ satisfies

$$\begin{aligned} \frac{2}{(\varepsilon\ell)^2} \partial_t \left[\mathcal{G}^2 \beta + \varepsilon (\mathcal{G}^2 \ell^2 U + U_z - U_{zz}) \right] & \quad (3a) \\ & = \frac{2}{\varepsilon\ell^2} e^z \partial_z \left[e^{-z} \alpha \partial_z (U - U^R) \right] \\ & \quad + k e^z \text{Im}\{\Psi^* \Psi_{zz}\} \end{aligned}$$

with boundary conditions

$$U(z=0) = U^R(z=0) = 10 \text{ m/s}$$

$$U_z(z=z_{\text{top}}) = U_z^R(z=z_{\text{top}}) = \gamma/1000$$

114 while the perturbation streamfunction amplitude $\Psi(z, t)$ satisfies

$$\begin{aligned}
 (\partial_t + ik\varepsilon U) & \left[-\mathcal{G}^2(k^2 + \ell^2) - \frac{1}{4} + \partial_z^2 \right] \Psi \\
 + ik\Psi & \left[\mathcal{G}^2\beta + \varepsilon(\mathcal{G}^2\ell^2 U + U_z - U_{zz}) \right] \\
 & = -\left(\partial_z - \frac{1}{2} \right) \left[\alpha \left(\partial_z + \frac{1}{2} \right) \Psi \right]
 \end{aligned} \tag{3b}$$

with boundary conditions

$$\Psi(z = 0) = \frac{gh}{f_0}$$

$$\Psi(z = z_{\text{top}}) = 0.$$

115 We have defined the nondimensional parameter $\mathcal{G}^2 := H^2 N^2 / (f_0^2 L^2)$, where f_0 is the coriolis
 116 parameter at 60°N, $N^2 = 4 \times 10^{-4}$ is the the stratification, and $L = 2.5 \times 10^5$ km is a horizontal
 117 length scale chosen to make non-dimensionalized U and Ψ variables have similar climatological
 118 variances. The linear relaxation towards $U^R(z) = 10 \text{ m/s} + (\gamma/1000)z$ on the right-hand side of
 119 Eq. (3a) is the force that maintains the typically strong polar vortex. Here $\gamma = 1.5 \text{ m s}^{-1} \text{ km}^{-1}$.
 120 The relaxation is mediated by a Newtonian cooling profile $\alpha(z)$, which is plotted in Fig. 1a, in its
 121 original dimensional units. Meanwhile, the lower boundary condition on Ψ comes from a bottom
 122 topography $h \cos(kx)$, where $h = 38.5$ m. This serves as a source of planetary waves.

123 There are two differences from Finkel et al. (2021), besides rearrangement. First, Finkel et al.
 124 (2021) had an erroneous but inconsequential negative sign in front of U_{zz}^R (their Eq. 3) which is
 125 corrected in Eq. (3a). Second, the left side of Eq. (3b) has two terms, $\pm ik\varepsilon \mathcal{G}^2 \ell^2 U \Psi$, which could
 126 be cancelled out; we have retained them both to maintain a term-by-term correspondence with the

127 original QGPV equation,

$$(\partial_t + \bar{u}\partial_x)q' + v'\partial_y\bar{q} = \text{sources} - \text{sinks}, \quad (4)$$

$$\text{where } q' = \nabla^2\psi' + \frac{f_0^2}{N^2}e^{z/H}\partial_z(e^{-z/H}\psi') \quad (5)$$

$$\text{and } v' = \partial_x\psi' \quad (6)$$

128 which will be important when deriving the enstrophy budget in section 5.

129 After discretizing to 27 vertical levels, we end up with a state space of dimension $d = 3 \times (27 - 2) =$
 130 75, with a state vector

$$\mathbf{X}(t) = \left[\text{Re}\{\Psi(t)\}, \text{Im}\{\Psi(t)\}, U(t) \right] \in \mathbb{R}^{75} \quad (7)$$

131 each of the three entries representing a vector with 25 discrete altitudes. We thus obtain a system
 132 of 75 ODEs, $\dot{\mathbf{X}}(t) = \mathbf{v}(\mathbf{X}(t))$. We furthermore perturb the system by stochastic forcing to represent
 133 unresolved processes such as smaller-scale Rossby and gravity waves, initial condition uncertainties,
 134 and sources of model error, an approach originally put forward by Birner and Williams (2008) and
 135 used more recently by Esler and Mester (2019). The forcing is white in time, giving an Itô diffusion

$$d\mathbf{X}(t) = \mathbf{v}(\mathbf{X}(t)) dt + \boldsymbol{\sigma}(\mathbf{X}(t)) d\mathbf{W}(t) \quad (8)$$

136 where $\mathbf{v}(\mathbf{x})$ (not to be confused with meridional wind velocity, v) is the drift function determined
 137 by Eqs. (3). $\mathbf{W}(t)$ is an $(m + 1)$ -dimensional white-noise process, and $\boldsymbol{\sigma} \in \mathbb{R}^{d \times (m+1)}$ is a matrix
 138 specifying the spatially smooth structure of the noise as Fourier modes in the vertical. $\boldsymbol{\sigma}$ could
 139 depend on the state vector \mathbf{X} , but for simplicity we fix it to a constant, defined as follows. At
 140 each timestep $\delta t = 0.005$ days, after incrementing the full system by $\delta\mathbf{X} = \mathbf{v}(\mathbf{X})\delta t$, we additionally
 141 increment the zonal wind profile by

$$\delta U(z) = \sigma_U \sum_{k=0}^m \eta_k \sin \left[\left(k + \frac{1}{2} \right) \pi \frac{z}{z_{\text{top}}} \right] \sqrt{\delta t} \quad (9)$$

142 where $\sigma_U = 1 \text{ m s}^{-1} \text{ day}^{-1/2}$, whose units reflect the quadratic variation of Brownian motion. The
 143 numerical scheme is known as Euler-Maruyama (see, e.g., Pavliotis 2014, ch. 5). Equation 9 fully
 144 defines the matrix σ . For $k = 0, \dots, m$, the k th column starts with 50 zeros, since there is no forcing
 145 on $\text{Re}\{\Psi\}$ or $\text{Im}\{\Psi\}$. The last 25 entries are evenly spaced samples of the sinusoidal factor in
 146 Eq. (9), all times σ_U .

147 The specific choice of stochastic forcing does affect the transition path statistics, but our method
 148 can be applied to any stochastic forcing. Because of the nonlinear coupling between $U(z)$ and $\Psi(z)$
 149 in Eqs. (3a) and (3b), the noise injected to U feeds to Ψ after a single timestep.

150 *b. Diagnostics*

151 Until section 5, we use two main diagnostics for visualization, the same as in Finkel et al. (2021).
 152 The first is zonal wind strength $U(z)$, an index for vortex strength which is used to define regimes
 153 A and B . The second is the meridional eddy heat flux $\overline{v'T'}(z)$, which quantifies the heat being
 154 advected into the polar region associated with the sudden warming, and in the quasi-geostrophic
 155 limit, the vertical propagation of Rossby waves. In the Holton-Mass model, this takes the form

$$\overline{v'T'}(z) = \frac{Hf_0}{R} \frac{\partial \overline{\psi'}}{\partial x} \frac{\partial \overline{\psi'}}{\partial z} \propto e^{z/H} |\Psi(z)|^2 \frac{\partial \varphi}{\partial z}, \quad (10)$$

156 where R is the ideal gas constant for dry air and φ is the phase of the complex-valued streamfunction
 157 Ψ . Hence the heat flux is related to the amplitude and phase tilt of the waves, both of which rise
 158 significantly during a SSW event. We also use the density-weighted vertical integral of heat flux,

$$\text{IHF}(z) := \int_0^z e^{-z'/H} \overline{v'T'}(z') dz' \quad (11)$$

159 which varies more smoothly than $\overline{v'T'}$ at any single altitude.

160 *c. Bistability*

161 We use the same constant parameters and boundary conditions as Finkel et al. (2021), which
162 give rise to two stable equilibria: a radiative equilibrium-like state, denoted **a**, and a disturbed state
163 **b**, in which upward propagating stationary waves flux momentum down to the lower boundary,
164 weakening zonal winds. Detailed bifurcation analysis by Yoden (1987a) and Christiansen (2000)
165 found a range of values for bottom topography h that create bistability. Figure 1(b,c) depicts
166 the zonal wind and streamfunction of these two equilibria. SSW events in this model are abrupt
167 transitions from the region near **a** to the region near **b**. If a strong wave from below happens
168 to catch the stratospheric vortex in a vulnerable configuration, then a burst of wave activity can
169 propagate upward, ripping apart the polar vortex and causing zonal wind to collapse (Charney
170 and Drazin 1961; Yoden 1987b). With certain parameters, the vortex can get stuck in repeated
171 “vacillation cycles”, in which the vortex begins to restore with the help of radiative forcing, only
172 to be undermined quickly by the wave. The situation of two well-separated equilibria is highly
173 idealized, and not a generic feature of climate phenomena; this system, with these parameters,
174 serves to demonstrate qualitative features of SSW, not represent the real stratosphere quantitatively.
175 Holton and Mass (1976); Yoden (1987b); Christiansen (2000), and Finkel et al. (2021) contain
176 further details.

177 A *transition path* is defined as an unbroken segment, or trajectory, of the system that begins in a
178 region A of state space (a neighborhood of **a**) and travels to another region B (a neighborhood of
179 **b**) without returning to A . As in Finkel et al. (2021), we define A and B based on the zonal-mean
180 zonal wind at $z = 30$ km:

$$A = \{\mathbf{x} \in \mathbb{R}^d : U(30 \text{ km})(\mathbf{x}) \geq 53.8 \text{ m/s}\} \quad (12a)$$

$$B = \{\mathbf{x} \in \mathbb{R}^d : U(30 \text{ km})(\mathbf{x}) \leq 1.75 \text{ m/s}\} \quad (12b)$$

181 where the velocity thresholds correspond to the vortex strength at 30 km for the fixed points \mathbf{a} and
182 \mathbf{b} , respectively.

183 An SSW event is then a transition from A to B , while the reverse, from B to A , represents the
184 recovery of the vortex. The definition of B modifies the widely used definition of Charlton and
185 Polvani (2007) in two ways. First, we use zonal wind at 30 km above the tropopause (in log-pressure
186 coordinates), because 30 km is where the zonal wind profile of \mathbf{b} reaches a minimum; Christiansen
187 (2000) used this same coordinate when studying the same model. (The standard 10 hPa pressure
188 level would correspond to $z = -7\text{km} \times \log(10/1000) - 10\text{km} \approx 22\text{ km}$ above the troposphere in
189 this model.) We also modify the zonal wind thresholds order to ensure that $\mathbf{a} \in A$ and $\mathbf{b} \in B$.

190 An important consequence of our A and B definitions is that the $A \rightarrow B$ transition path takes
191 ~ 80 days. By design, this includes the slow initial *preconditioning* stage of vortex breakdown in
192 advance of the ~ 10 -day time horizon that traditionally comprises an SSW event. In this paper,
193 ‘SSW event’ should be interpreted as both the preconditioning and the ensuing vortex collapse.

194 Figure 2 shows timeseries of U and $\overline{v'T'}$ at several different altitudes as the system goes through
195 several transition paths in a long simulation. As in Fig. 2 of Finkel et al. (2021), orange strips denote
196 $A \rightarrow B$ transitions while green strips denote $B \rightarrow A$ transitions. The long periods in between, which
197 we call the $A \rightarrow A$ and $B \rightarrow B$ phases, demonstrate the bistable nature of regimes A and B . The
198 fleeting $A \rightarrow B$ phase, however, is what we seek to understand. When the system is en route from
199 A to B , we say it is *(AB)-reactive*, using a term from chemistry literature where the passage
200 from A (reactant) to B (product) models a chemical reaction. The following section will introduce
201 the *reactive density* $\pi_{AB}(\mathbf{x})$ and associated *reactive current* $\mathbf{J}_{AB}(\mathbf{x})$ which help us visualize the
202 transition as a path distribution through state space and make the foregoing observations more
203 quantitative.

3. The reactive density and reactive current: a distribution over transition paths

We consider the long-time behavior of our stochastic Holton-Mass model $\mathbf{X}(t)$ undergoing transitions between states A and B . Aggregating together statistics from only the transition paths yields a probability distribution called the *reactive density* $\pi_{AB}(\mathbf{x})$, defined such that

$$\pi_{AB}(\mathbf{x}) d\mathbf{x} = \mathbb{P}\{\mathbf{X}(t) \in d\mathbf{x} | \mathbf{X}(t) \text{ is in transition from } A \text{ to } B\} \quad (13)$$

where $d\mathbf{x}$ is a small region about \mathbf{x} . One could estimate π_{AB} by binning samples from a long simulation, but including only those samples in transit directly from A to B . Associated to π_{AB} is a vector field called the *reactive current* $\mathbf{J}_{AB}(\mathbf{x})$, which quantifies the probability flux passing through \mathbf{x} per unit time only during transition paths. Roughly speaking, π_{AB} specifies where transition paths go, and \mathbf{J}_{AB} specifies how they move. Below we define them formally, but Fig. 3(a-c) gives some intuition by projecting them on the subspace (U, IHF) at $z = 10, 20, \text{ and } 30$ km. Background shading indicates the strength of π_{AB} , and arrows indicate the magnitude and direction of \mathbf{J}_{AB} . Overlaid in thin blue lines are ten randomly sampled transition paths from the long ergodic simulation. These sample paths cluster in the same regions of state space identified as high-probability under π_{AB} , and on average flow along the arrows, corroborating qualitatively that $\pi_{AB}(\mathbf{x})$ and \mathbf{J}_{AB} describe the location and evolution of the model in state space.

The transition path ensemble shows marked differences between altitudes. At $z = 10$ km, the vortex strength (U) of states \mathbf{a} and \mathbf{b} is about the same, but the IHF is very distinct. The reactive current aligns with the IHF axis. Mathematically, this reflects the lower boundary condition $U(z = 0) = U^R(z = 0)$. Physically, this means that the heat flux due to the wave is the dominant physical process, with only small changes in zonal wind strength. The higher altitude of $z = 30$ km, by contrast, exhibits a large reduction in zonal wind strength, but only in the late stages of the process.

225 In fact, the pattern of reactive density π_{AB} at $z = 30$ km (panel c) tells us that this final deceleration
226 is quite sudden: the magnitude of π_{AB} is large near A , meaning transition paths linger there for a
227 long time and only slowly crawl downward and to the right. But at the point $\text{IHF}(30 \text{ km}) \approx 2.5 \times 10^4$
228 $\text{K}\cdot\text{m/s}$, $U(30 \text{ km}) \approx 30 \text{ m/s}$ (the region marked by a dotted circle in panels c and f), π_{AB} reduces in
229 magnitude and the reactive current spreads out widely as it turns downward toward set B . This is a
230 signal that the transition paths are becoming both faster and more variable.

231 As a further point of comparison with \mathbf{J}_{AB} , we have plotted the minimum-action pathway from
232 A to B with thick cyan lines (section 3 of the supplement specifies the numerical method). This
233 represents the most likely transition path in the low-noise limit (e.g., Freidlin and Wentzell 1970;
234 E et al. 2004; Forgoston and Moore 2018), and indeed it follows the direction of reactive current.
235 With finite noise, however, the transition path ensemble spreads significantly around the minimum-
236 action pathway, especially at the higher altitude of 30 km in the late stage of the transition process.
237 Because of this, it is not possible for *any* single pathway, minimum-action or not, to meaningfully
238 represent the full ensemble.

239 We will show that the slow, initial phase of SSW involves *preconditioning* of the vortex: gradual
240 erosion of the wind field by the stochastic forcing into a configuration that is especially susceptible
241 to wave propagation. Once the wave burst is triggered, it imparts swift changes to the entire
242 zonal wind profile. However, the bulk of SSW progress, probabilistically speaking, occurs in the
243 preconditioning phase. Below we make this qualitative description precise by relating the reactive
244 current to the forecast functions from Finkel et al. (2021): the committor and expected lead time
245 metrics.

246 *a. Mathematical relationship between current, committor, density, and rate*

247 To formalize the description above and interpret the current rigorously, some definitions are
248 in order, including a brief recap of the quantities from Finkel et al. (2021). Let us fix an initial
249 condition $\mathbf{X}(t_0) = \mathbf{x}$ with a vortex that is neither strong nor fully broken down, so $\mathbf{x} \notin A \cup B$. $\mathbf{X}(t)$
250 will soon evolve into either A or B , since both are attractive. The probability of hitting B first is
251 called the *forward committor* (to B):

$$q_B^+(\mathbf{x}) = \mathbb{P}_{\mathbf{x}}\{\mathbf{X}(\tau_{A \cup B}^+(t_0)) \in B\} \quad (14)$$

252 where the subscript \mathbf{x} denotes a conditional probability given $\mathbf{X}(t_0) = \mathbf{x}$, and $\tau_S^+(t_0)$ is the *first*
253 *hitting time* after t_0 to a set $S \subset \mathbb{R}^d$:

$$\tau_S^+(t_0) = \min\{t > t_0 : \mathbf{X}(t) \in S\}. \quad (15)$$

254 Like the expected lead time introduced below, the committor (under various aliases) predates TPT
255 as an object of interest in the study of rare events (Du et al. 1998; Bolhuis et al. 2002). However,
256 as we will see below, it is a key ingredient in any TPT analysis.

257 Our system is autonomous, with no external time-dependent forcing, so we can set $t_0 = 0$ and drop
258 the argument from $\tau_{A \cup B}^+$ without loss of generality. The autonomous assumption can be relaxed,
259 either by augmenting \mathbf{x} with a periodic variable for time (e.g., to include the seasonal cycle) or by
260 augmenting A and B to include initial and terminal times (e.g., to better examine climate change
261 effects). Periodic- and finite-time TPT has been presented formally in Helfmann et al. (2020), and
262 we have applied it to a dataset of state-of-the-art ensemble forecasts in Finkel et al. (2022). As
263 a conceptual demonstration, however, the autonomous Holton-Mass model makes for a clearer
264 exposition.

265 While $\tau_{A \cup B}^+$ itself is a random variable, one can take its expectation to obtain the *expected lead*
 266 *time* (to B),

$$\eta_B^+(\mathbf{x}) := \mathbb{E}_{\mathbf{x}}[\tau_{A \cup B}^+ | \tau_B^+ < \tau_A^+], \quad (16)$$

267 in other words, the expected time of arrival to B conditional on hitting B first. Finkel et al. (2021)
 268 described q_B^+ and η_B^+ in detail, as they are central quantities for forecasting, and visualized them in
 269 their Figs. 2c,d and 3c in the observable subspace (U , IHF). We do the same here, but additionally
 270 we overlay the reactive current. In Fig. 3(d,e,f), background shading represents the expected lead
 271 time and black contours represent committor level sets of 0.1, 0.2, 0.5, 0.8, and 0.9.

272 The committor's contour structure differs a lot between altitude levels. At 10 and 30 km (panels d
 273 and f), the contours have kinks. Depending on the initial condition, either a fluctuation in U or IHF
 274 might have a greater effect on the committor. The intermediate altitude of 10 km seems special in
 275 having committor contours that align with the IHF axis along the main channel of reactive current.
 276 In other words, $q_B^+(\mathbf{x})$ is well-approximated by a linear function of $U(20 \text{ km})$, which is consistent
 277 with the finding in Finkel et al. (2021) that the 21.5-km altitude holds the most predictive power
 278 for q_B^+ .

279 \mathbf{J}_{AB} is related to q_B^+ , generally flowing up the committor gradient. But \mathbf{J}_{AB} contains some key
 280 information that the committor does not. As a *forecast* function, the committor does not distinguish
 281 $A \rightarrow B$ transitions from $B \rightarrow B$ transitions, where the system leaves state B (beginning to recover),
 282 but then falls back to the weak-vortex state. To isolate the transition events from A to B , we need
 283 to introduce the *backward committor* (to A):

$$q_A^-(\mathbf{x}) = \mathbb{P}_{\mathbf{x}}\{\mathbf{X}(\tau_{A \cup B}^-(t_0)) \in A\} \quad (17)$$

284 where $\tau_S^-(t_0)$ is the *most recent hitting time*

$$\tau_S^-(t_0) = \max\{t < t_0 : \mathbf{X}(t) \in S\} \quad (18)$$

285 Intuitively, $q_A^-(\mathbf{x})$ is the probability of the system at point \mathbf{x} last came from A , not B . The backward-
 286 in-time probabilities refer specifically to the process $\mathbf{X}(t)$ *in steady-state*, allowing us once again
 287 to set $t_0 = 0$. In other words, $q_A^-(\mathbf{x})$ depends explicitly on the *steady-state probability density* $\pi(\mathbf{x})$,
 288 where $\pi(\mathbf{x}) d\mathbf{x} = \mathbb{P}\{\mathbf{X}(t) \in d\mathbf{x}\}$ is the long-term (climatological) probability of finding the system
 289 in a small region $d\mathbf{x}$ about \mathbf{x} .

290 Having defined both forward and backward committors, we can express the reactive density as

$$\pi_{AB}(\mathbf{x}) = \frac{1}{Z_{AB}} \pi(\mathbf{x}) q_A^-(\mathbf{x}) q_B^+(\mathbf{x}) \quad (19)$$

291 where Z_{AB} is a normalizing constant such that the right-hand side integrates to one. The associated
 292 reactive current can in turn be expressed

$$\mathbf{J}_{AB}(\mathbf{x}) = q_A^- q_B^+ [\pi \mathbf{v} - \nabla \cdot (\mathbf{D}\pi)] \quad (20)$$

$$+ \pi \mathbf{D} [q_A^- \nabla q_B^+ - q_B^+ \nabla q_A^-], \quad (21)$$

293 where the diffusion matrix $\mathbf{D}(\mathbf{x}) = \frac{1}{2} \boldsymbol{\sigma}(\mathbf{x}) \boldsymbol{\sigma}(\mathbf{x})^\top$, and ∇ represents the gradient operator over state
 294 space.

295 Eq. (21) is a specific expression for the current of a diffusion process of the form (8), which is
 296 the same general formulation as our model. But a more illuminating and general definition is its
 297 connection to the *rate*, or inverse return time, of the event (approximately (1700 days)⁻¹ for the
 298 Holton-Mass model with our chosen parameters). Let C be a closed hypersurface in \mathbb{R}^d which
 299 encloses A and is disjoint with B ; we call this a *dividing surface*. In the context of the diagrams in
 300 Fig. 3, C is any curve separating region A from region B . Then we have

$$\oint_C \mathbf{J}_{AB} \cdot \mathbf{n} dS = \text{Transition rate} \quad (22)$$

301 where \mathbf{n} is an outward unit normal from C and dS is a surface area element. The integral rela-
302 tionship (22) holds for any dividing surface, implying that the current is divergence-free outside
303 of A and B , but has a source in A and a sink in B (see Vanden-Eijnden (2006) for a thorough
304 mathematical explanation of \mathbf{J}_{AB} .) This constraint immediately implies a link between magnitude
305 and width of \mathbf{J}_{AB} streamlines. In Fig. 3(c,f), the strong magnitude of \mathbf{J}_{AB} near \mathbf{a} implies a thin
306 central channel, and strict constraints on the mechanisms of early SSW onset. In other words, the
307 initial preconditioning phase can only happen in a small number of ways. On the other hand, the
308 subsequent weakening of \mathbf{J}_{AB} between $q_B^+ = 0.5$ and $q_B^+ = 0.8$ (in the boxed region of Fig. 3c,f)
309 implies that paths fan out across state space, becoming more variable. This spreading, or diversity
310 of events, is only with respect to U and IHF at 30 km; at the lower altitudes, the current remains
311 strong and narrow all the way through the transition process (Fig. 3, columns 1 and 2).

312 The reactive current and density characterize the transition path ensemble across the continuum
313 of possible pathways, providing more information than the numerical value of the rate itself. Given
314 any user-defined set of coordinates, the reactive current projection maps the transition paths in
315 those coordinates, as a statistical ensemble with average behavior and variability. Below, following
316 a brief note on the computational method, sections 4 and 5 demonstrate how to use reactive current
317 and density to describe climatology and strengthen physical understanding of a rare transition
318 event.

319 *b. Computational method*

320 The quantities presented in section 3, as well as the results to follow, could be computed directly
321 by running a model for long enough to undergo a large number of SSW events and analyzing the
322 statistics of those transitions. This procedure, which we call the “ergodic simulation” (ES) method,
323 is possible in the 75-dimensional Holton-Mass model, and we have performed such a simulation

324 of 10^6 days for validation purposes. However, this can be a major computational barrier in global
325 climate models when the numerical integration is costly and the return period is long compared
326 to the simulation timestep. Anticipating the need for fundamentally different techniques in high-
327 dimensional state spaces, we have instead used the dynamical Galerkin approximation (DGA;
328 Thiede et al. 2019; Strahan et al. 2021). A large collection of trajectories are launched in parallel
329 with initial conditions distributed across state space, each one running for only a short time relative
330 to the return period. Here we use 3×10^5 trajectories of length 20 days each, which is shorter than
331 the 80-day duration of a single SSW event and much shorter than the 1700-day return period.
332 Afterward, we assemble all these pieces together to estimate the quantities of interest, exploiting
333 the Markov property. The total simulation time is not always reduced by this method—in our case,
334 the short simulations total 6×10^6 days compared with the 1×10^6 -day ES—but the format opens
335 the door for many interesting possibilities, such as massive parallelization and adaptive sampling.
336 In particular, as we show in Finkel et al. (2022), DGA is uniquely positioned to exploit large
337 ensembles of short weather forecasts from high-fidelity operational models.

338 The basic DGA algorithm for rare event analysis has been described and tested in a recent series
339 of articles (Thiede et al. 2019; Strahan et al. 2021; Finkel et al. 2021; Antoszewski et al. 2021).
340 It is closely related to the “analogue Markov chain” approach of Lucente et al. (2021). Recently,
341 an approach to learning neural network approximations of forecast functions using short trajectory
342 data was introduced in Strahan et al. (2022). Due to the dependence on steady state and backward-
343 in-time quantities, a full TPT analysis as carried out in this paper requires additional calculations
344 beyond what is described in Finkel et al. (2021). We leave these details to the supplement in order
345 to keep the focus on the results of our TPT analysis, which are robust with respect to algorithmic
346 parameters.

4. SSW composites

Here we explain the traditional notion of a rare event ‘composite’ and contrast it with the composite intrinsically defined by TPT. The results are qualitatively similar, but the TPT description allows a rigorous mathematical connection to the reactive current and SSW rate.

The standard “composite” of an SSW event is a day-by-day aggregate of all the SSW events in a given dataset, aligned by the central warming date. This can include statistics, such as the mean and quantiles, of any observable function, such as the zonal-mean zonal wind or heat flux. Charlton and Polvani (2007) and Charlton et al. (2007) used this method to describe SSW climatology and establish benchmarks for stratosphere-resolving GCMs. We form a standard composite of $U(30$ km) from our Holton-Mass model in Fig. 4a, averaging together 300 events from a long ergodic simulation.

Here, we propose a complementary “TPT composite” based on reactive density. Instead of aligning events by the central warming date, we align the events by a general coordinate $f(\mathbf{x})$, which can be user-defined but must fulfill the minimal criterion of increasing from A to B , so it represents some objective notion of progress. At any progress level f_0 , the TPT composite is defined by restricting the reactive density $\pi_{AB}(\mathbf{x})$ to the level set $\{\mathbf{x} : f(\mathbf{x}) = f_0\}$. Fixing $f = f_0$ is not the same as fixing the lead time τ_B^+ , because the threshold might be crossed at different times by different transition paths. Note that $f(\mathbf{x})$ is a deterministic function of initial condition \mathbf{x} , unlike the hitting time τ_B^+ , which is a random variable that changes between realizations launched from the same initial condition. Therefore, τ_B^+ cannot itself be used as a progress coordinate.

In Fig. 4b,c, we juxtapose alternative composites with the standard warming date coordinate $-\tau_B^+$. In panel b, we aggregate paths based on the negative expected lead time $-\eta_B^+$ defined above: the *expected* time until the central warming date. $-\eta_B^+$ is the deterministic progress function that

370 is closest (in the mean-square sense) to the random progress function $t - \tau_B^+$ defining traditional
 371 composites. Panel c uses an altogether different progress metric, the committor q_B^+ itself, which
 372 increases from 0 on A to 1 on B .

373 The traditional and TPT composites are similar in shape, with an initially gradual decay in
 374 $U(30 \text{ km})$ accelerating into a rapid decline in the final few days. As a function of $-\eta_B^+$, $U(30 \text{ km})$
 375 accelerates steadily through the whole transition, in both the traditional and TPT composites. But
 376 as a function of committor, $U(30 \text{ km})$ decreases linearly at first and then accelerates downward
 377 between $q_B^+ = 0.6$ and $q_B^+ = 0.7$. According to the standard composite, $U(30 \text{ km})$ becomes steadily
 378 less variable over time, with the whole ensemble collapsing into a single path by construction, as
 379 $t = 0$ is the time of the event when $U(30 \text{ km}) = 0$. But when viewed as a function of expected lead
 380 time or committor, $U(30 \text{ km})$ becomes more variable in the middle of the path, starting at $\eta_B^+ \approx 50$
 381 days or $q_B^+ \approx 0.65$ and lasting until the end, when $\eta_B^+ \rightarrow 0$ and $q^+ \rightarrow 1$.

382 The same variability is reflected in Fig. 3c,f. In the boxed region, the reactive density weakens
 383 and the reactive current spreads out, some paths turning straight downward into B and others
 384 accumulating still more heat flux before making the plunge. The q_B^+ and η_B^+ contours in Fig. 3f
 385 convey geometrically how it is possible to have such wide variation in zonal wind strength even
 386 at a fixed expected lead time. Along the central channel of strong reactive current, where most of
 387 the transition paths flow, the committor and expected lead time have an approximately (negative)
 388 linear relationship. But in the weak- U flank of the current, especially in the boxed region, the q_B^+
 389 level sets “unkink” to align with the IHF axis while the η_B^+ level sets turn downward to align with
 390 the U axis. The lowest visible level set of η_B^+ thus spans a range of vortex strengths of $U(30 \text{ km})$.

391 Physically, the TPT composites are more variable than the traditional composite because $-\eta_B^+$,
 392 the expected lead time—a deterministic function—is a coarser description than $t - \tau_B^+$, a random
 393 variable. The former is an average over all realizations, while the latter takes on a specific value for

394 each realization, which is not actually known until after the warming occurs. Given only information
395 on the resolved variables $\Psi(z, t)$ and $U(z, t)$ at a given time, the TPT composite is the best one can
396 do. The expected lead time quantifies SSW predictability, as established in Finkel et al. (2021).
397 Here, we additionally incorporate the backward committor q_A^- via the reactive density π_{AB} , and so
398 restrict focus to *transition* events—“major warmings”—from A to B .

399 As a loose analogy, a student’s progress toward a degree can be measured objectively in course
400 credits. On the other hand, first-year exams might weed out half of all students, which means that the
401 *probabilistic* half-way point usually comes before half of required credits are done. A third metric,
402 the time until graduation, can vary due to random effects like gap years and pandemics, which
403 can cause a student to space their course load unevenly in time. Each cross-section of the student
404 population—conditioning on a fixed number of credits completed, probability of graduation, or
405 expected time until graduation—is a different statistical ensemble, each one conveying different
406 information.

407 Going forward, we will use the committor as the progress coordinate of choice. That way,
408 each point along the composite is an average over trajectories that are equally predictable in their
409 probability to reach B , i.e., to proceed to an SSW. Often it is not just a singular coin toss that
410 determines the fate of $\mathbf{X}(t)$, but a whole sequence of ‘coin tosses’—random turns through state
411 space—aligning in just such a way to navigate from A to B . With the committor as a progress
412 coordinate, the ‘coin tosses’ are equidistributed along the horizontal axis, though they may not be
413 equidistributed in time.

414 The same composite technique can be used to visualize the vertical wind structure at different
415 stages. Fig. 5 plots $U(z)$ and $\overline{v'T'}(z)$ as altitude-indexed probability distributions at a series of
416 committor level sets: $q_B^+ = 0.1, 0.5, \text{ and } 0.9$. The widening variability with increasing committor
417 is faintly visible at low altitudes, but increases dramatically above ~ 23 km, where at the $q_B^+ = 0.9$

418 level, the mean state (orange curve) falls well below the median state (central gray envelope.) This
 419 means the distribution of transition states is skewed left by a minority of paths with early collapse
 420 of upper-level winds. At the same committor range of 0.5-0.9, the vertical profile of meridional heat
 421 flux inflates dramatically. The altitude range of $z = 20\text{-}25$ km is the key transition region, below
 422 which zonal wind evolves relatively smoothly and with a symmetric distribution, and above which
 423 it varies rapidly with a skewed distribution. $\overline{v'T'}(z)$ is maximum near this altitude. We speculate
 424 that the underlying reason is the Newtonian cooling profile $\alpha(z)$, which has its own transition
 425 region centered at 25 km. It is not surprising that zonal wind just below, at 21.5 km, is an optimal
 426 linear predictor, as we found in Finkel et al. (2021).

427 **5. A wave-mean flow interaction perspective**

428 The previous section presented \mathbf{J}_{AB} and π_{AB} as functions of two basic observables, zonal wind and
 429 integrated heat flux, and constructed a composite evolution of these observables. In this section, we
 430 incorporate more detailed physical knowledge to improve the interpretability of our TPT results. In
 431 particular, we manipulate the dynamical equations to derive an enstrophy budget in the Holton-
 432 Mass model, which reveals a more natural set of coordinates that separates conservative from
 433 non-conservative processes. By visualizing the current in these coordinates, we identify physical
 434 drivers of each stage in the transition process. Our goal is twofold: first, to show how TPT can be
 435 formulated for any observables, and second, more narrowly in the context of this study, how the
 436 dynamics become more clear when those observables are well-chosen.

437 A common diagnostic for wave-mean flow interaction systems is the wave activity, $\mathcal{A} =$
 438 $\overline{\rho_s q'^2} / (2\partial_y \bar{q})$, whose evolution is related to the Eliassen-Palm (EP) flux divergence (Andrews
 439 and McIntyre 1976). Yoden (1987b) used wave activity extensively to analyze the vacillating
 440 regime (our set B) of the Holton-Mass model, in particular the upward wave propagation that

441 destabilizes the vortex. Below we derive a related set of equations for the eddy enstrophy, which
 442 enjoys a simpler balance equation and which we have found is better numerically suited for TPT
 443 analysis.

444 The first step in deriving the EP relation is to multiply the QGPV equation (4) by q' and take a
 445 zonal average, yielding

$$\partial_t \left(\frac{\overline{q'^2}}{2} \right) + \overline{v'q' \partial_y \bar{q}} = \overline{q'(\text{sources} - \text{sinks})} \quad (23)$$

446 We wish to work with the projected version of the equation, Eq. (3b), rather than the original
 447 PDE, to account for the approximation $\sin^2(\ell y) \approx \varepsilon \sin(\ell y)$ introduced by Holton and Mass (1976)
 448 for projecting quadratic nonlinearities. The procedure is summarized below, and spelled out more
 449 thoroughly in section 4 of the supplement.

450 Because of the ansatz (2), q' is represented in the projected equations by

$$\begin{aligned} q' &\longleftrightarrow \left[-\mathcal{G}^2(k^2 + \ell^2) - \frac{1}{4} + \partial_z^2 \right] \Psi \\ &=: (-\delta + \partial_z^2) \Psi \end{aligned} \quad (24)$$

451 where \longleftrightarrow denotes correspondence between the full governing equations and the projected, non-
 452 dimensionalized equations in the Holton-Mass model. Recall that Ψ is the complex amplitude for
 453 the zonal-perturbation streamfunction $\psi'(x, y, z, t)$, in geostrophic balance with the wind (u, v) .

454 As a general rule, the zonal average of the product of two wave quantities ψ'_1 and ψ'_2 of the form
 455 in Eq. (2).is found by the following formula:

$$\begin{aligned} \overline{\psi'_1 \psi'_2} &= \overline{\text{Re}\{\Psi_1 e^{ikx}\} \text{Re}\{\Psi_2 e^{ikx}\}} \\ &= \text{Re}\{\Psi_1^* \Psi_2\} \end{aligned} \quad (25)$$

456 Therefore, we multiply both sides of Eq. (3b) by the complex conjugate of (24) and take the real
 457 part to obtain

$$\partial_t \mathcal{E} + F_q \beta_e = D \quad (26a)$$

458 where

$$\begin{aligned} \mathcal{E} &= \frac{1}{2} e^z |(-\delta + \partial_z^2) \Psi|^2 \\ &\longleftrightarrow \frac{1}{2} \overline{q'^2} \end{aligned} \quad (26b)$$

459 represents the eddy enstrophy;

$$\begin{aligned} F_q &= k e^z \text{Im}\{\Psi^* \Psi_{zz}\} \\ &\longleftrightarrow \overline{v' q'} \end{aligned} \quad (26c)$$

460 represents the meridional eddy PV flux;

$$\begin{aligned} \beta_e &= \mathcal{G}^2 \beta + \varepsilon (\mathcal{G}^2 \ell^2 U + U_z - U_{zz}) \\ &\longleftrightarrow \partial_y \overline{q} \end{aligned} \quad (26d)$$

461 represents the meridional PV gradient; and

$$\begin{aligned} D &= -\text{Re} \left\{ e^z [(-\delta + \partial_z^2) \Psi^*] \times \right. \\ &\quad \left. \left(\partial_z - \frac{1}{2} \right) \left[\alpha \left(\partial_z + \frac{1}{2} \right) \Psi \right] \right\} \\ &\longleftrightarrow \overline{q'(\text{sources} - \text{sinks})} \end{aligned}$$

462 represents the production and dissipation of enstrophy.

463 The standard EP relation would be found by dividing both sides by the meridional PV gradient
464 β_e , as in Yoden (1987b). Instead, we next turn to the mean-flow equation (3a), which is an evolution
465 equation for the PV gradient β_e rather than U directly. Multiplying through by β_e , we find

$$\partial_t \Gamma = R \beta_e + F_q \beta_e \quad (27a)$$

466 where

$$\Gamma := \left(\frac{\beta_e}{\varepsilon \ell} \right)^2 \quad (27b)$$

$$R := \frac{2}{\varepsilon \ell^2} e^z \partial_z \left[e^{-z} \alpha \partial_z (U - U_R) \right] \quad (27c)$$

467 The new quantity $\Gamma(z)$ is the squared meridional gradient of zonal-mean potential vorticity, which
468 is highly correlated to zonal wind strength $U(z)$ in the Holton-Mass model. R is a relaxation
469 coefficient for Γ , strengthening the vortex via radiative cooling.

470 The advantage of this alternative EP relation is now clear: adding together Eqs. (26) and (27),
471 the meridional PV transport $F_q \beta_e$ cancels to give

$$\partial_t (\Gamma + \mathcal{E}) = R \beta_e + D. \quad (28)$$

472 In this form, all the dissipative effects are contained on the right-hand side via the cooling coefficient
473 $\alpha(z)$, which appears both in D and R . $\Gamma + \mathcal{E}$ would be conserved, at every altitude separately, in
474 the absence of dissipation and stochastic forcing. In this limit, an increase in eddy enstrophy
475 \mathcal{E} can only occur at the expense of the mean PV gradient characterized by Γ . Of course, both
476 non-conservative effects—dissipation and stochastic forcing—are critically important; vacillation
477 cycles and transitions are possible only because the Holton-Mass model, like the full atmosphere,
478 is an open system. The utility of Eq. (28) is to isolate those nonconservative effects as almost
479 extrinsic inputs.

480 *a. The importance of non-conservative effects as a function of altitude, inferred from reactive*
481 *current*

482 Dissipation and forcing act to disrupt the conservation of $\Gamma + \mathcal{E}$, with a specific pattern shown
483 in Fig. 6. The reactive current is shown at three altitudes, as in Fig. 3, but this time in the space

484 $(\Gamma^{1/2}, \mathcal{E}^{1/2})$ instead of (U, IHF) . We take square roots because the visualizations are more clear,
 485 and the units of s^{-1} are more comparable with those of zonal wind $U(z)$ and radiative cooling
 486 $\alpha(z)$. (Note that the fixed point \mathbf{b} in panel (d) appears to have committor < 1 ; this is possible
 487 when projecting out nonlinear coordinates because set B is defined based on the 30-km level,
 488 and the state-space regions that resemble \mathbf{b} at 10 km may not resemble it at 30 km.) In the upper
 489 stratosphere, at $z = 30$ km (panels c and f), the main channel of reactive current flows along a circular
 490 arc, approximately conserving $\Gamma + \mathcal{E}$, all the way through the $q_B^+ = 0.9$ surface: the evolution of an
 491 SSW is a nearly conservative interaction between waves and the mean flow right up to the end.
 492 Then, the current weakens in magnitude and spreads out, indicating the critical non-conservative
 493 processes at the end, where the breaking and dissipation of the anomalous waves cements the SSW
 494 event. Just as in the (U, IHF) space, the reactive density π_{AB} decreases along that circular arc,
 495 meaning the transition paths accelerate.

496 On the other hand, \mathbf{J}_{AB} projected at $z = 10$ km (panels a and d) shows that the dynamics are never
 497 conservative in the lower stratosphere: the initial motion points not along a circular arc but directly
 498 leftward, such that $\Gamma + \mathcal{E}$ is decreasing from the start. From the enstrophy budget (28), we conclude
 499 that a combination of dissipation and stochastic forcing acts strongly at 10 km to precondition the
 500 vortex. The next subsection shows that stochastic forcing plays the more decisive role.

501 Finally, consider the middle altitude of 20 km, where \mathbf{J}_{AB} has a shape that is intermediate between
 502 the current at 10 and 30 km. It does not have distinctly positive or negative curvature, but flows
 503 along a straight channel from A to B . 20 km seems to be in just the right altitude range to feel
 504 significant dissipation and stochastic forcing—a feature of the lower boundary—but also to channel
 505 a good share of the loss of Γ to the gain of \mathcal{E} , a quasi-conservative property of the loftier 30 km. The
 506 resulting committor, expected lead time, and reactive current are approximately linear functions of

507 $\Gamma^{1/2}(20 \text{ km})$ and $\mathcal{E}^{1/2}(20 \text{ km})$. Indeed, the wind and heat flux at 20 km were the most useful for
508 prediction in (Finkel et al. 2021, their section 4).

509 Fig. 7a,b,c show the composite evolution of $\Gamma + \mathcal{E}$ in orange, along with Γ in blue and \mathcal{E} in pink,
510 at the same three altitudes 10, 20, and 30 km. All three altitudes show evidence of dissipation, with
511 $\Gamma + \mathcal{E}$ weakening as the committor increases, but with distinct differences in the rates. The $\Gamma + \mathcal{E}$
512 composite is concave up at 10 km, implying dissipation is most important at the early stage, when
513 the predictability of the event is limited. At 20 km, the composite is practically linear, implying
514 that dissipation maintains a constant role in the event’s evolution, gradually opening the valve to
515 wave propagation at the last stage of the event. At 30 km, the composite is concave down: the
516 flow is initially conservative, with exchange between mean flow and eddies at the onset of vortex
517 breakdown, followed by strong dissipation of the waves when the event is all but assured.

518 At 20 and 30 km, the distribution of $\Gamma + \mathcal{E}$ begins symmetric, with the mean (orange) tracking
519 the median (near the center of the dark gray band). Then between $q_B^+ = 0.6$ and 0.7, the lower tail
520 of the distribution expands quickly, skewing the distribution negative. The distribution at 10 km
521 maintains a slight negative skew for the entire transition path. The skewness reflects the occurrence
522 of “minor warmings” preceding the SSW, when the vortex begins to break down, but partially
523 recovers before the final event.

524 The composites, as well as the reactive currents, support the notion of the “typical” transition
525 path as an initially non-conservative creep at low altitudes, opening up a valve to allow waves to
526 propagate upward, finally yielding a very abrupt collapse at high altitudes follows after a long,
527 mostly conservative phase. With the enstrophy budget (28), we can assess the importance of each
528 term by plotting those composites as well. Fig. 7d,e,f show the composite evolution of each term at
529 each altitude: $R\beta_e$ (the relaxation of the squared mean PV gradient, Γ) in blue, D (the dissipation
530 of enstrophy, \mathcal{E}) in pink, and $\beta_e F_q$ (the transfer of enstrophy from Γ to \mathcal{E}) in black, all normalized

531 by the total $\Gamma + \mathcal{E}$ at each level to account for the altitude-dependent differences in variability.
 532 This allows us to compare how strong each dissipative force is *relative* to the total budget. The
 533 sum $(R\beta_e + D)/(\Gamma + \mathcal{E})$ —the normalized, deterministic tendency $\partial_t(\Gamma + \mathcal{E})/(\Gamma + \mathcal{E})$ —is shown as
 534 a dashed orange curve. Note that this tendency is positive at 10 and 20 km even though $\Gamma + \mathcal{E}$
 535 is actually decreasing. Without stochastic forcing, the system will always approach state **a** or **b**,
 536 depending on where the initial condition falls relative to the surface dividing the two attractors.

537 To quantify the critical role of stochastic forcing in effecting the transition at each committor
 538 level, we define the stochastic tendency of $\Gamma + \mathcal{E}$ along transition paths:

$$\begin{aligned} \mathcal{L}_{AB}[\Gamma + \mathcal{E}](\mathbf{x}) = & \tag{29} \\ & \lim_{\Delta t \rightarrow 0} \mathbb{E} \left[\frac{(\Gamma + \mathcal{E})(\mathbf{X}(t + \Delta t)) - (\Gamma + \mathcal{E})(\mathbf{X}(t - \Delta t))}{2\Delta t} \right. \\ & \left. \middle| \mathbf{X}(t) = \mathbf{x} \text{ and } \mathbf{X}(t) \text{ is in transition} \right] \tag{30} \end{aligned}$$

539 which is related to the ordinary infinitesimal generator \mathcal{L} (see Oksendal (2003) for mathematical
 540 background and the appendix of Finkel et al. (2021) for its application to the Holton-Mass model).
 541 The supplement describes the numerical procedure to approximate \mathcal{L}_{AB} using short trajectories
 542 and with a finite lag time. There, we show that $\mathcal{L}_{AB}f(\mathbf{x})$ is related to $\mathbf{J}_{AB} \cdot \nabla f(\mathbf{x})$ for any observable
 543 f , so it is appropriate to view the arrows in Fig. 3 and 6 as a proxy for the stochastic tendencies of
 544 the projected observables.

545 We introduce \mathcal{L}_{AB} to compare with the deterministic tendency $\partial_t(\Gamma + \mathcal{E})(\mathbf{x})$, which for a diffusion
 546 process of the form (8) is simply $\mathbf{v}(\mathbf{x}) \cdot \nabla(\Gamma + \mathcal{E})(\mathbf{x})$ by the chain rule. Their difference shows the
 547 impact of stochastic forcing responsible for transitions. More specifically, $\mathcal{L}_{AB} - \partial_t$ averaged over
 548 a committor level q_0 highlights the stochastic effects responsible for taking the system from q_0 to

549 $q_0 + dq$. Often it is not just a single coin flip that decides the fate of $\mathbf{X}(t)$, but a whole sequence of
550 random turns through state space aligning in just such a way to navigate from A to B .

551 The role of stochasticity is most stark at 10 and 20 km (panels (d) and (e)) and for $q_B^+ < 0.5$,
552 where $\mathcal{L}_{AB}(\Gamma + \mathcal{E})$ is negative while $\partial_t(\Gamma + \mathcal{E})$ is positive, due to a strong positive tug of radiative
553 cooling versus the weak dissipation of enstrophy. As q_B^+ increases, the stochastic and deterministic
554 tendencies grow closer together: the more likely the transition to B , the easier it is for deterministic
555 drift to carry it out alone. At 30 km (panel f), all forms of dissipation and forcing start out *relatively*
556 small compared to the magnitude of $\Gamma + \mathcal{E}$, but as the path progresses they all diverge away from
557 zero. Most notably, the stochastic and deterministic tendencies never diverge very far; if anything,
558 stochastic noise *slows* the collapse of $U(30 \text{ km})$ at the end. It seems that to achieve the $A \rightarrow B$
559 transition, which is defined entirely in terms of $U(30 \text{ km})$, the most common mechanism is a
560 persistent negative push applied to lower altitudes, and this ultimately sets up the higher altitudes
561 for more sudden, deterministic collapse after the “hard work” of eroding the vortex from below is
562 mostly finished.

563 In summary, the TPT diagnostics have demonstrated that the SSW process begins with steady,
564 significant decay of the PV gradient (and its squared gradient, Γ) at lower altitudes, driven by the
565 stochastic forcing, with only conservative changes taking place at higher altitudes. This precondi-
566 tioning of the vortex opens up a valve to the mid-stratosphere. In the late stages of the transition,
567 starting between $q_B^+ = 0.6$ and 0.7 , the upper-level winds decline very suddenly. This begins con-
568 servatively as eddies grow, exchanging energy with the mean flow, and finishes non-conservatively,
569 as friction dissipates the waves.

570 **6. Conclusion**

571 Transition path theory (TPT) is a mathematical framework that can be used to assess the near-
572 term predictability and long-term climatology of anomalous weather events. The framework lends
573 itself naturally to events associated with regime transitions, but in general it can be applied to
574 more general anomalies. The key is to be able to define a suitable reaction coordinate linking
575 the event to the mean state. We have analyzed the statistical ensemble of sudden stratospheric
576 warmings (SSWs) in the Holton-Mass model. Probability densities and currents tell us how the
577 system evolves through state space during the vortex breakdown. The reactive current allows us
578 to condition dynamical tendencies on the occurrence of the rare transition event. By overlaying
579 \mathbf{J}_{AB} over observable subspaces at different altitudes in the stratosphere, we have identified the key
580 roles of dissipation and stochastic forcing patterns that bring about transition paths. The stochastic
581 driving represents the effects of unresolved Rossby and gravity waves that have been stripped from
582 this highly truncated model. These non-conservative processes, stochastic driving in particular,
583 seem to act most forcefully at lower altitudes early in the transition process, conditioning the
584 vortex, while the higher altitudes are shielded from significant dissipation. It is only late in the
585 transition process, after the likelihood of the event has surpassed 60%, that the upper-level winds
586 play a significant role in the dynamics.

587 This work is an early application of TPT to atmospheric science. We believe it holds potential
588 as a framework for forecasting, risk analysis, and uncertainty quantification. Thus far, it has been
589 used mainly to analyze protein folding in molecular dynamics, but is now being applied in diverse
590 fields such as social science (Helfmann et al. 2021), as well as ocean and atmospheric science
591 (Finkel et al. 2020; Helfmann et al. 2020; Miron et al. 2021, 2022). TPT results are best interpreted
592 when viewed in a physically meaningful observable subspace. With the rather simple Holton-

593 Mass model, we identified such a subspace based on an enstrophy budget. In different versions of
594 quasigeostrophic dynamics, the wave activity (Nakamura and Solomon 2010; Lubis et al. 2018)
595 and other diagnostics based on the transformed-Eulerian-mean (Andrews and McIntyre 1976) are
596 likely to be informative coordinates.

597 Significant challenges remain for deploying TPT analysis at scale to state-of-the-art climate
598 models. We have used the DGA short trajectory analysis algorithm to compute TPT quantities.
599 One important limitation of this computational pipeline is the data generation step. We used a
600 long ergodic trajectory to sample the attractor, which served the double purpose of seeding initial
601 data points for short trajectories and providing a ground truth for validating the accuracy of DGA.
602 In some cases, short trajectory data already exist, e.g., from the subseasonal-to-seasonal (S2S)
603 database (Vitart and Robertson 2018), which we have used recently in (Finkel et al. 2022) to
604 estimate centennial-scale SSW rates from only 21 years of ensemble forecasts. In other cases, it is
605 advantageous to generate fresh data in undersampled regions of state space, which would require
606 more advanced sampling methods such as the adaptive sampling strategies proposed in Lucente
607 et al. (2021) and Strahan et al. (2022) or rare event simulation schemes such as in Mohamad and
608 Sapsis (2018); Ragone et al. (2018); Webber et al. (2019); and Ragone and Bouchet (2020).

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623 *Data availability statement.* The code to produce the data set and results, either on the Holton-
624 Mass model or on other systems, is publicly available at [https://github.com/justinfocus12/](https://github.com/justinfocus12/SHORT)
625 SHORT. Interested users are encouraged to contact J.F. for more guidance on usage of the code.

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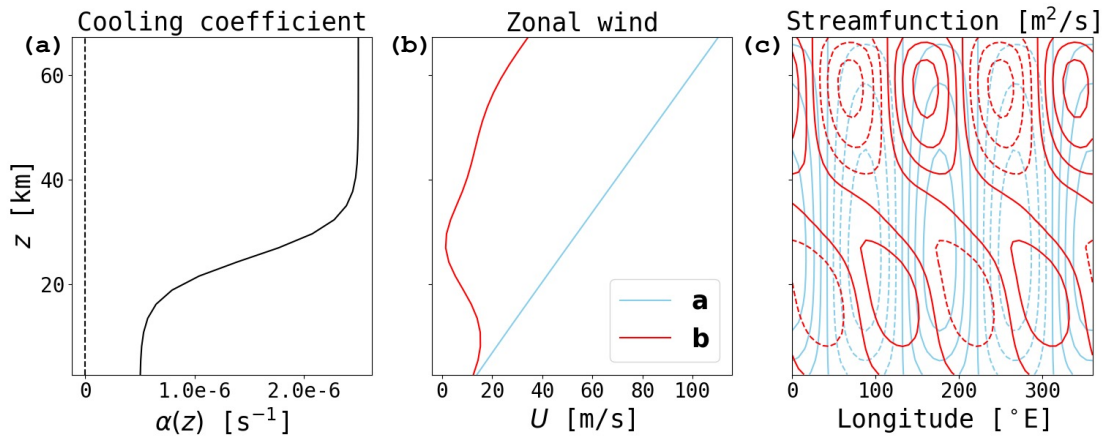
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LIST OF FIGURES

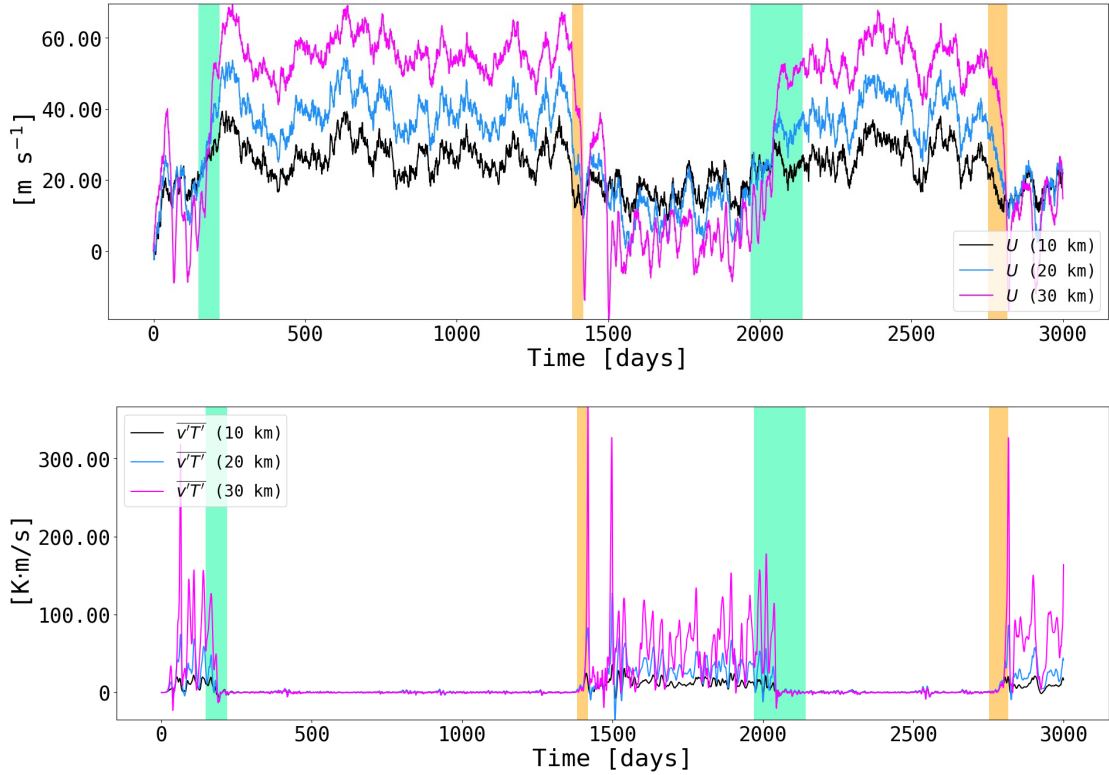
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- 814 **Fig. 3. Currents, densities, committors, and expected lead times.** (a): Background shading is the reactive density π_{AB} , on a log scale. Thin blue lines are ten randomly selected transition paths from the long control simulation. Thick cyan curve is the minimum-action path from A to B . Also overlaid is a vector field representing reactive current \mathbf{J}_{AB} . The subspace is (U, IHF) evaluated at $z = 10$ km. Positions of the fixed points **a** and **b** are marked. Arrows represent \mathbf{J}_{AB} . (b, c): Same as (a), but at $z = 20$ and 30 km respectively. (d) The expected lead time η_B^+ is shaded as background color, and level sets of the committor q_B^+ 0.1, 0.2, 0.5, 0.8, and 0.9 are overlaid as black curves. (e, f): Same as (d), but at $z = 20$ km and 30 km respectively. A box marks a transition region between narrow, constrained current and wide, dispersed current. See text for a description. 45
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- 824 **Fig. 4. Composites evolution of SSW events.** Orange curves plot the mean value of $U(30$ km) at a given stage in the transition process; expanding gray envelopes show the middle 25-, 50-, and 90-percentile ranges. We use three different notions of progress: hitting time to B ($t - \tau_B^+$, panel a), expected hitting time to B ($-\eta_B$, panel b), and committor (q_B^+ , panel c). 46
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- 827
- 828 **Fig. 5. Vertical profiles of transition states and tendencies.** Left column: $U(z)$ averaged over $q_B^+ = 0.1, 0.5$, and 0.9. Orange curve is the mean, and gray envelopes represent the middle 25-, 50-, and 90-percentile ranges. Dashed blue and red curves represent $U(z)$ for the fixed points **a** and **b**. Right column: same as left, but for eddy meridional heat flux $\overline{v'T'}$ 47
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- 832 **Fig. 6. Current in wave-mean flow coordinates.** Same as Fig. 3, but for a different observable subspace $(\Gamma^{1/2}, \mathcal{E}^{1/2})$ instead of (U, IHF) . See text for definitions. Eddies are characterized by RMS perturbation PV, $\mathcal{E}^{1/2}$, and the mean flow by the zonal mean PV gradient, $\Gamma^{1/2}$ 48
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- 835 **Fig. 7. Enstrophy budget analysis through the $A \rightarrow B$ transition.** (a) Blue, pink, and orange curves represent mean values of Γ , \mathcal{E} , and their sum at $z = 10$ km, conditioned on the system being in a transition path and near a given committor level (which varies along the horizontal axis). Gray envelopes represent the middle 25, 50, and 90-percentile ranges of $\Gamma + \mathcal{E}$; when the orange curve is not at the center of the gray envelopes, the distribution is skewed. (b, c): same as (a), but at $z = 20$ and 30 km respectively. (d) Solid orange curve shows the expected tendency of $\Gamma + \mathcal{E}$ at 10 km, again conditioned on being in a transition path and near a given committor level. Dashed orange curve shows the deterministic tendency at the same committor levels; the difference between the two indicates the role of stochastic forcing. Blue curve shows the relaxation of Γ (the squared meridional PV gradient), pink curve shows the dissipation of enstrophy, and black curve shows the meridional transport of PV, $F_q \beta_e$, which when negative indicates a gain for \mathcal{E} at the expense of Γ . The sum of the blue and pink curves gives the dashed orange curve. (e, f): same as (d), but at $z = 10$ and 20 km respectively. All
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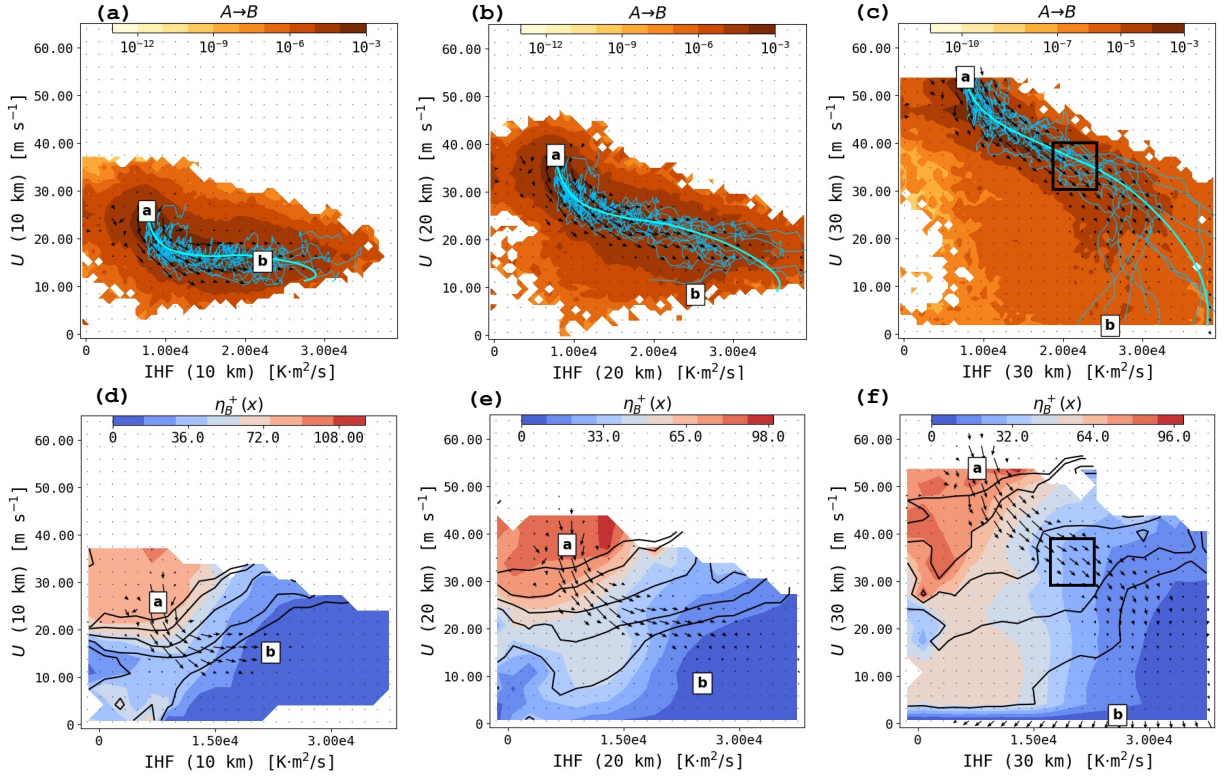
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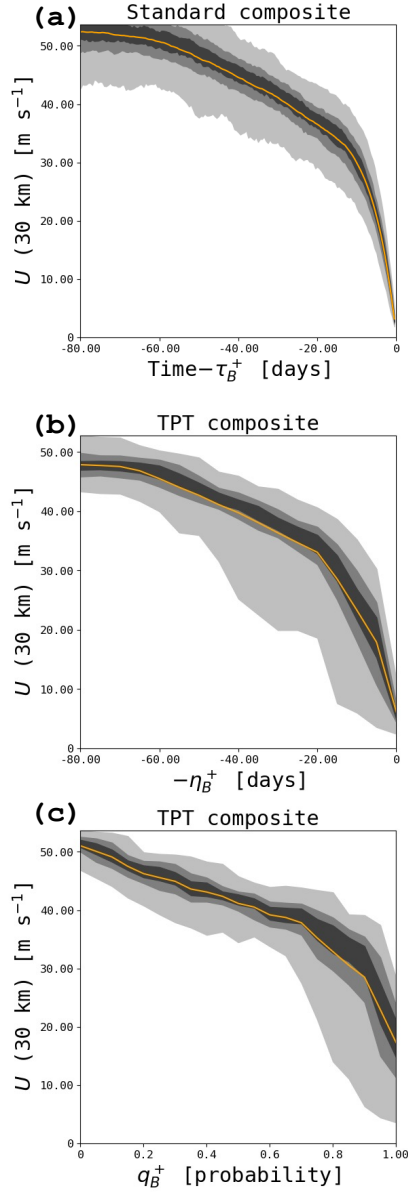
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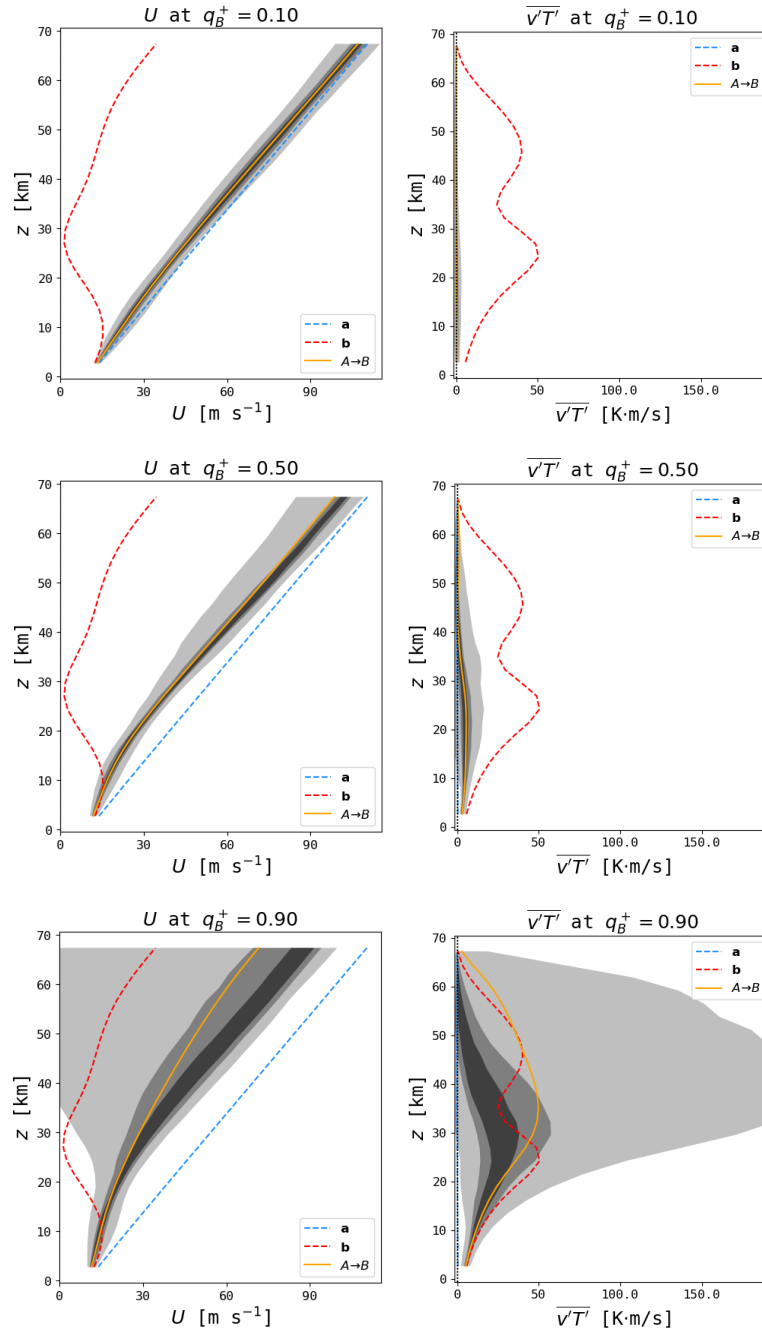
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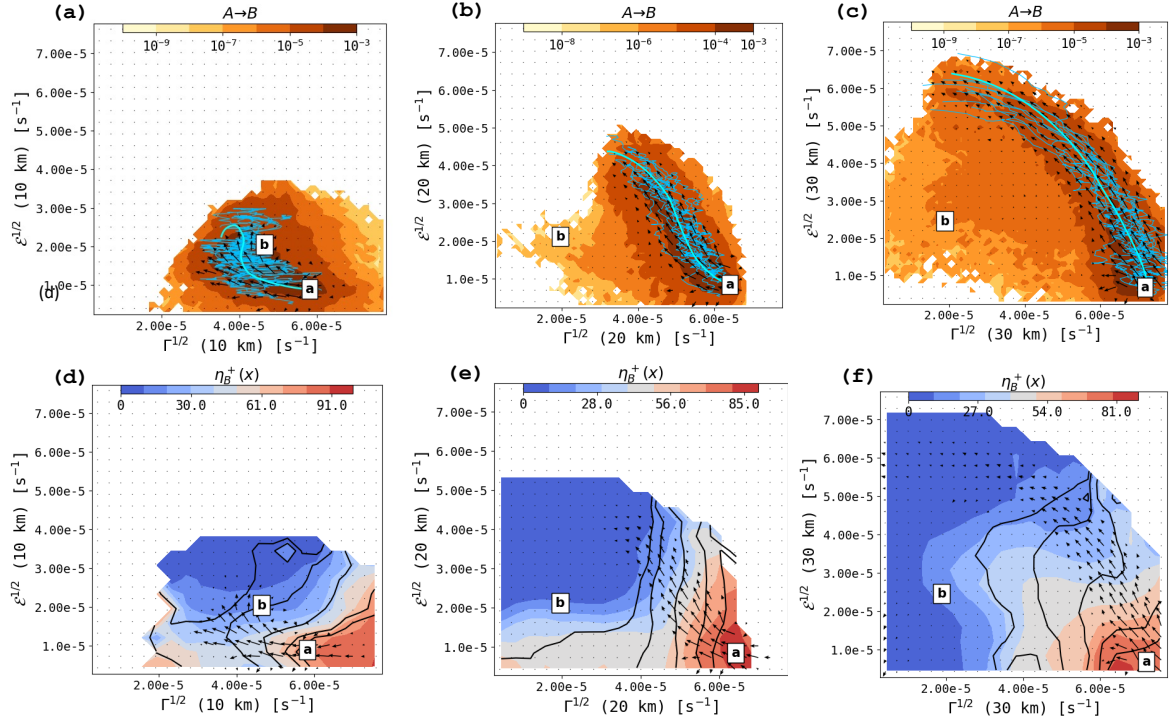
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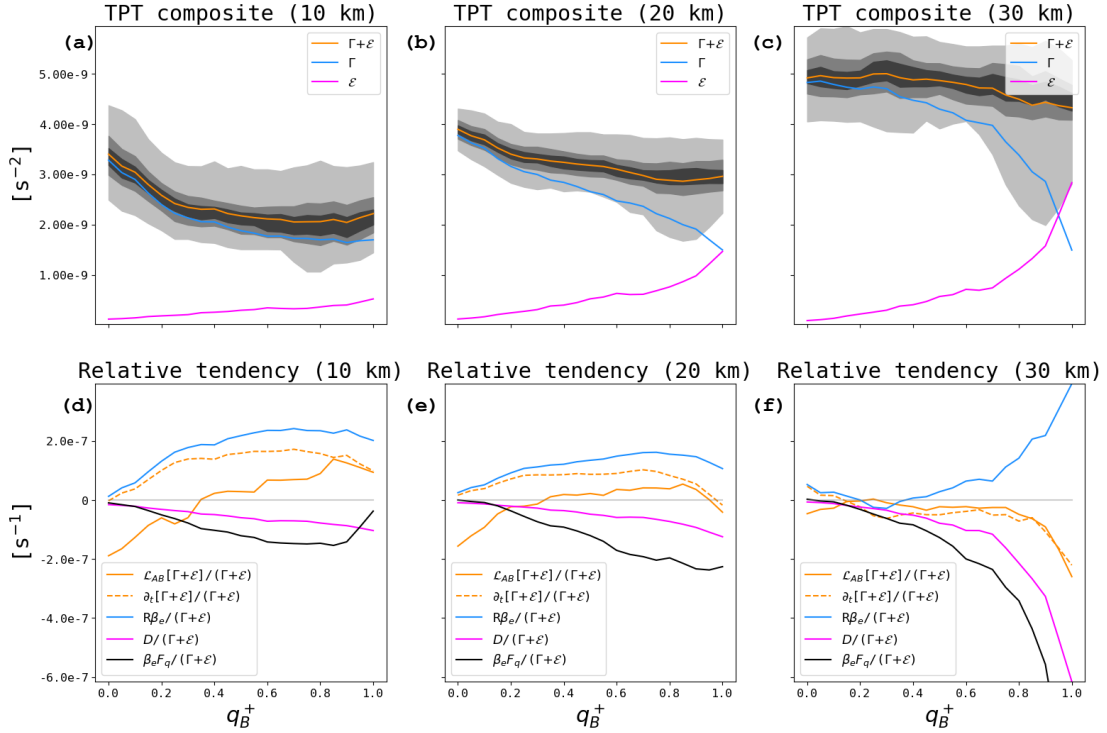
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